PHOTOMETRY OF THE EARTH FROM MARINER II

Robert L. Wildey

Division of Geological Sciences and Mount Wilson and Palomar Observatories, California Institute of Technology and Carnegie Institution of Washington

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ABSTRACT

The Earth tracking system aboard the Mariner II spacecraft has collected photometric observations of the Earth as a by-product of its navigational duties on the flight to Venus. The observations show good agreement with the phase curve for the Earth that was previously found by observing the earth-lit moon. Diurnal variations in brightness are correlated with the fraction of the Earth's disk covered by land.

Introduction. All previous photometric studies of the total visual flux (i.e. integrated over the apparent disk) of the earth have been made by measuring the brightness of that portion of the moon which is illuminated only by sunlight that has been reflected from the earth. The classical work in this field has been that of Danjon (1954), using the "cat-eye" photometer, and is the source of the presently accepted value of the albedo of the earth. Such measurements are based on a visual nulling of the brightness-suppressed sunlit portion of the moon against the unsuppressed earthlit portion. The amount of suppression, accurately known, constitutes the fundamental datum. It is then necessary to make a completely independent set of comparisons of the sunlit moon with a constant source, usually the sun, before an absolute visual flux for the earth can be obtained. Danjon's method solves quite well the problem of atmospheric extinction.

There are three principal difficulties to be encountered with the use of the moon as an intermediary in doing photometry of the earth. (I) The sunlit portion is the source of considerable scattered light in the earthlit image which may make measurements systematically too high, and as one approaches full moon the amount of scattered light increases while the amount of earth-shine decreases. (2) To within rather narrow limits, the fixed location of an observatory imposes an unwanted correlation between the earth's phase angle, as seen from the moon, and the geographic longitude and latitude of the sublunar point. The latitude bias thus introduced in the earth's phase curve can be investigated by examing the observations for seasonal variations. Although the longitude effects can be investigated by making measurements from observatories well distributed over the earth's surface, such observations have been previously collected only in France. A restrictive corollary of this effect is that the earth's phase angle cannot be held constant while the geography presented upon its disk is varied and any implied brightness changes are measured. (3) The lunar neighborhoods on opposite sides of the terminator that are intercompared may possess intrinsic sources of

brightness difference which would be present even under illumination by the same flux. They may differ slightly in normal albedo, and must also possess differences, of varying magnitude, due to the non-degeneracy of the moon's photometric function in any of its three degrees of freedom.

Such difficulties are alleviated when one does photometry of the earth from space.

Such observations are the useful by-product of the Mariner - 11 spacecraft's guidance instrumentation.

Observations. The reduceable data were collected during the 52 day period from September 29 to November 22, 1962. Prior to the earlier date the telemetry indicated an abnormally low signal by nearly a factor of 100, the explanation for which remains elusive. Recovery was sudden and was followed by indications of normal operation.

Later than November 22, the temperature, which had veen rising rapidly, was too high to be corrected for in the data reduction. It will be seen that the method of correction for the temperature dependence of responsivity begins to fail toward the latter part of the 52 day period.

The photometer used to collect these observations constitutes one aspect of the the Long Range Earth Sensor (LRES). This device has been described in detail by G. W. Meisenholder in the internal publications of the Jet Propulsion Laboratory (SPS 37-6, -9, -14, and -16, Vol. II) and by McLauchlan (1964). Inasmuch as these reports are of limited edition, a short redescription will be given.

The LRES uses a single, end-on, 3/4 -in. diameter, S-11 photocathode Dumont photomultiplier tube whose field is limited by a fine aperture stop. Both stop and detector are behind the focal plane. Coincident at all times with the focal plane is a modulating mask attached to a 22 cps vibrating read. Ahead of the modulating mask is a fixed aperture, preceded in the optical train by the objective, which is a 7 - element f1.2 lens of 2 inch focal length. The purpose of the LRES is to provide information regarding orientation in space for the purpose of performing navigational and orientational

maneuvers. In this function the photomultiplier tube serves only as a light detector and has no position discrimination by itself.

The vibrating reed moves through a sufficient arc so that the modulating mask completely uncovers the fixed aperture at the extremes of motion. The photoelectric output, therefore, is a series of pulses. The shapes of the knife edges of the modulating mask are such that a shift of the position of the Earth image in hinge causes a variation in the output pulse width, while a shift in roll causes a variation in phase (or time) of the photoelectric output relative to the reed motion. These properties of the output waveform are detected and provide dc error signals for attitude correction. Also present in the output are signals which indicate that an object is being tracked by the sensor and the amount of light being sensed.

The threshold of the unit is approximately 5×10^{-11} W/CM 2 (blackbody bolometric at 6000° K) and the maximum permissible sustained signal is 5×10^{-7} W/CM 2 .

The photometer signal was calibrated, relatively, over the entire dynamic range to be encountered on the mission, by use of an earth simulator. Absolute calibration was obtained with the aid of a National Bureau of Standards lamp. No in-flight calibration existed. The foot-candle (ft-cd) has been chosen as the unit in which to express the results of the photometry. It is a standard unit and is based on a radiation bandpass that does not differ from the photometer bandpass by as much as any other logical choice except, of course, a mean monochromatic flux corresponding to the effective wavelength of the photometer. It can be readily shown that the reduction equation is exactly given by:

$$S_{\bullet} = \frac{\left\langle f_{\bullet}(\lambda) \right\rangle_{V} \left\langle f_{o}(\lambda) \right\rangle_{SII}}{\left\langle f_{o}(\lambda) \right\rangle_{SII} \left\langle f_{o}(\lambda) \right\rangle_{V}} \cdot S_{o} \cdot R (T, 70^{\circ}) \cdot \frac{I_{\bullet}}{I_{o}}$$
 (I)

where the subscript V following the expectation-value brackets denotes that the quantity within the brackets has been weighted over wavelength according to the response function of the eye (Walsh, 1953). A subscript SII implies that the weighting function has been

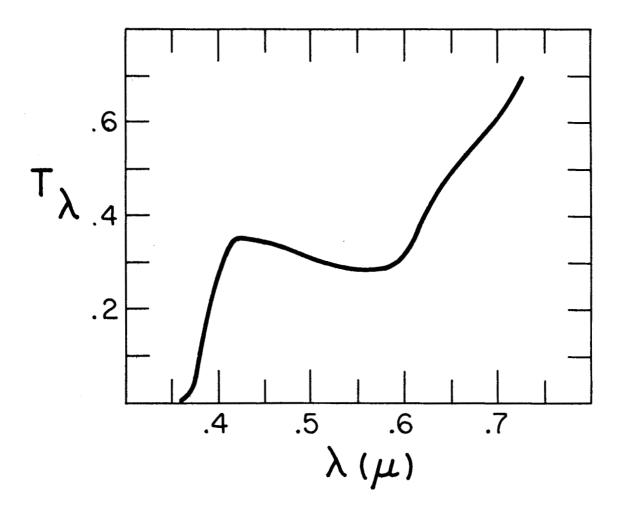


Fig. 1. Spectral transmission curve for the objective lens of the type used in Mariner II's Long Range Earth Sensor.

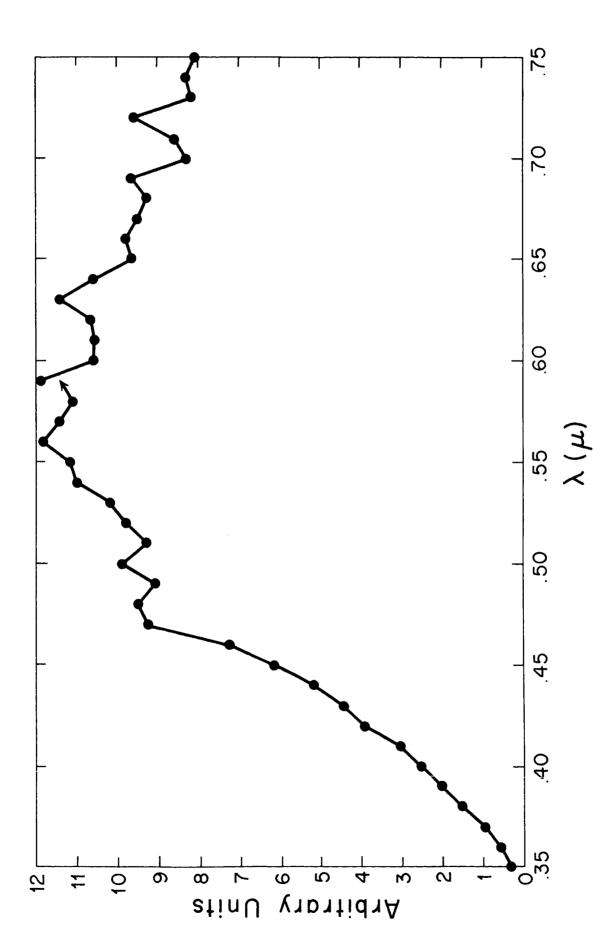


Fig. 2. Spectral energy distribution of calibration source.

multiplied by the transmission of the objective lens (Figure 1) the response of an SII photoemissive surface (Radio Corporation of America, 1958).

The quantities $f_o(\lambda)$ and $f_o(\lambda)$ are the spectral energy distributions, in arbitrary units, of the standard source and the earth flux, respectively. S_o is the foot-candle output of the standard source. $R(T, 70^o)$ is the ratio of the response at 70^oF , at which temperature the laboratory measurements were made, to the response at T, the temperature of the photometer at a given measuring time, which is also in the telemetry record.

I is the photoelectric current produced by the earth flux and I_o is the photo-current produced by the standard source.

Measurements of $f_o(\lambda)$ were undertaken in the laboratory and are shown in Figure 2. $f_o(\lambda)$ was taken as the product of three factors. The first is the empirecal solar continuum published by Minnaert (1953). The second is one minus the Frauhofer line blanketing coefficient published by Michard (1950). The third term must allow for the wavelength dependence of the overall reflectivity of the earth. The only information bearing upon this are Danjon's 3 color observations. Danjon gives a color index (C.1.) for earthshine transformed to the magnitude-color system of Rougier (1937). He finds a seasonal variation. The median color has been adopted, which should be an optimum procedure since the sub-Mariner point was always close to the equator. A heuristic procedure was adopted to make use of this C.1. of 0.68. Rougier's C.1. of the moon shining by direct sunlight is 1.10. Thus the earth's reflectivity at the blue effective wavelength is -0.42 magnitudes, or 47.3 percent, higher than it is at the visual effective wavelength. It remains to specify the effective wavelengths. For this purpose an ordinary color equation was assumed connecting C.1. to the Johnson (1955) BV system .

$$B - V = A (C.I.) + C$$
 (2)

Using Rougier's values of C.1. for the sun (0.79) and the moon, and values of B - V for the sun of 0.60 (Stebbins and Kron, 1957) and for the moon of 0.85 (van den Bergh, 1962; Wildey and Pohn, 1964) equation 2 was solved for A. Interpreting A in

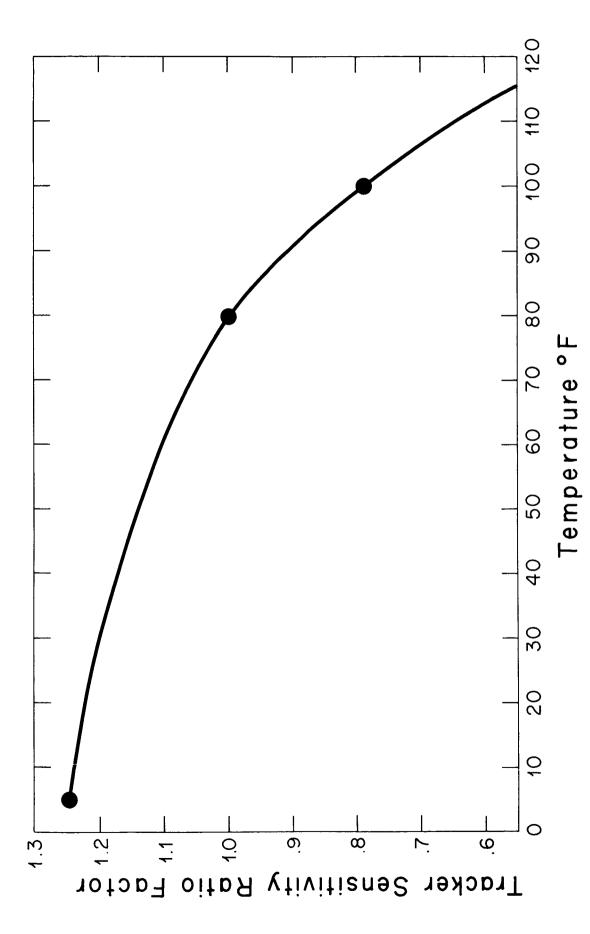


Fig. 3. Relative response, in arbitrary units of radiative power, as a function of temperature .

terms of the wavelength baselines in the usual way (e.g. <u>Wildey and Murray</u>, 1964; <u>Mathews and Sandage</u>, 1963) and assuming the effective wavelengths of B to be λ 4430A and of both V and Rougier's yellow magnitude to be λ 5540A, the blue effective wavelength of Rougier becomes λ 4230A. The curve of the reflectivity of the earth versus wavelength was then assumed to be linear with a value of 1.47 at λ 4230A and 1.00 at λ 5540A. This source of uncertainty may imply a systematic error in the photometry of 10 or 20 percent.

Three data points were obtained in the laboratory relating photometer responsivity to temperature. They are plotted in Figure 2. A smooth curve, whose functional form could not be specified, was fitted to these three points by eye. The accuracy of the correction provided implies it is an insignificant source of systematic error for most of the data of the present study. Barring a highly improbable interpretation of the later data, however, the temperature renders the latter part of the observations unacceptable for the study of other than diurnal brightness variations.

The possibility of the existence of other unknown sources of systematic error cannot be completely discounted in view of the unexplained behavior of the LRES signal at the beginning of the flight.

The only important source of random error is the number of significant digits carried in the telemetry and results in an uncertainty of about ± 2 percent. Extracting a tabulation of observations from the telemetry record becomes primarily a problem of identifying the recorded times when a given signal is holding fast and when the frequency of truncation between two consecutive signal numbers is equal.

The observations are tabulated in Table 1. Most of the column headings are self-explanatory. Column one lists the geocentric julian day beyond 2437937. Column 7 lists the luminous flux of the earth normalized to a distance of one astronomical unit assuming an inverse square dependence. This flux has been temperature corrected.

Results. The integral brightness versus phase of the earth over the range in phase

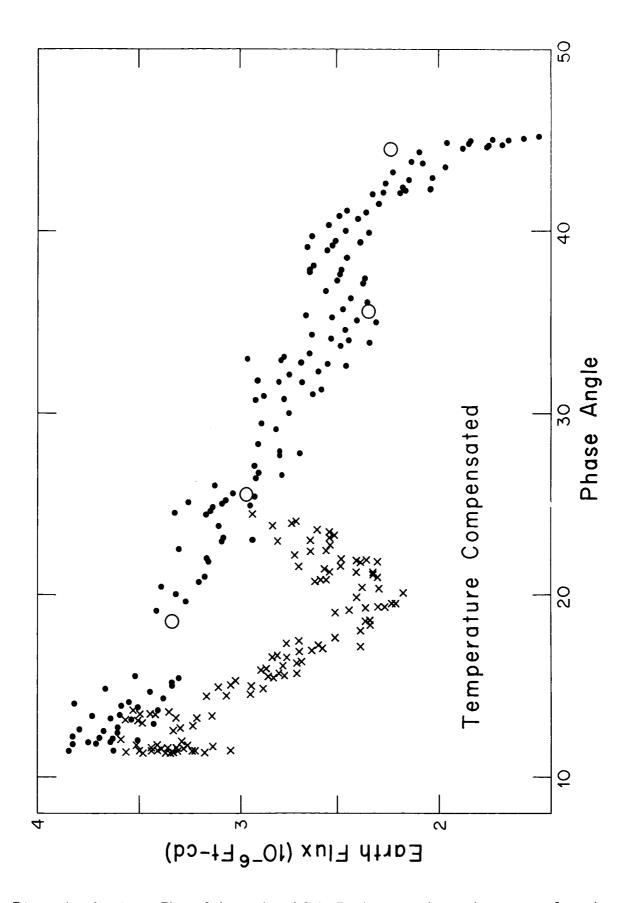


Fig. 4. Luminous Flux of the entire visible Earth versus the angle, as seen from the Earth, between the sun and the direction of measurement. Plot has been corrected for temperature variations of the photometer.

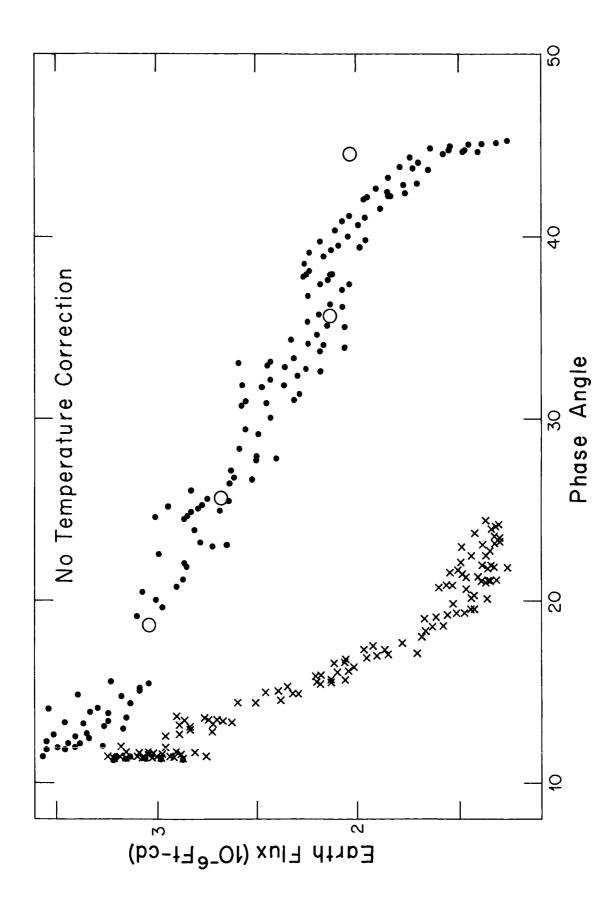


Fig. 5. As in Figure 4 except that no correction has been made for the temperature variation between laboratory and celestial environment.

angle possible for the present study is shown in Figure 3. Figure 4. Figure 4. Shows, for comparison, the same plot uncorrected for temperature.

Chronologically the phase—angle decreases from the maximum 45° shown to the minimum indicated in the figure and then again increases. This minimum corresponds to about the time that the temperature begins to rise sharply. The failure to reproduce the phase curve recorded before minimum phase, as the Sun-Earth-Mariner angle is now increased, is apparently a failure of Figure 2 to represent the responsivity variation at these higher temperatures.

The scatter present in the sequences of Figures X and X is merely the diurnal brightness variation whose periodicity is very short compared to the time required to produce a significant change in phase-angle.

It can be seen in Figures 2 and 2 that there is also present a semi-periodic light variation superimposed on the diurnal effect, of approximately the same or slightly larger amplitude, but with a characteristic period of about 5 to 6 days. No satisfactory explanation for this phenomenon can be offered at present.

The large open circles represent Danjon's mean curve for the Earth's integral brightness versus phase, to which has been applied a normalization scale factor to fit the present data at Danjon's minimum phase angle. Danjon's relation has the seasonal variation averaged out. It should thus correspond to the Mariner photometry which was always near zero geographic latitude, except for the longitude effect mentioned earlier. The agreement between the indirect and the direct photometry is good. Unfortunately, investigation at the larger more critical phase angles was not permitted by the trajectory. Hopefully, this important data will be provided by the LRES data collected on the Mariner B flight to Mars.

The scale factor found necessary to be applied to Danjon's data in order to produce agreement with the present measurements was 0.482. In view of the size of previously investigated systematic errors and especially because the abnormal

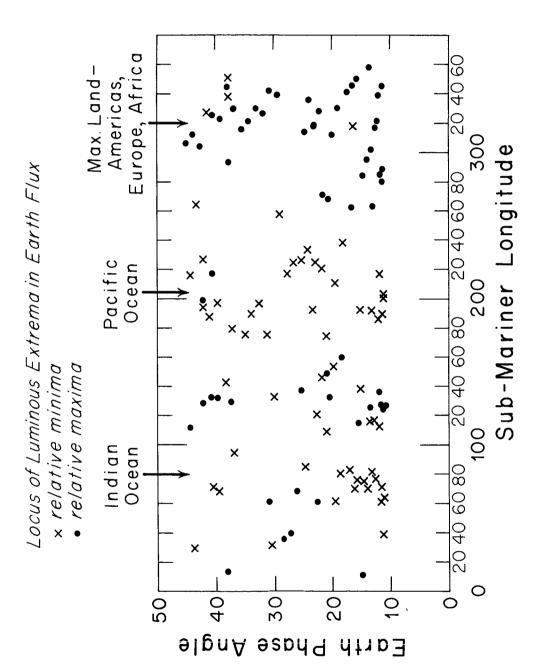


Fig. 6. Plot of the phase angle of the Earth versus the geographic longitude of the sub-Mariner point for those times when the Long Range Earth Sensor is recording a relative maximum or minimum in signal.

behavior of LRES during the early period of the flight has not been satisfactorily explained, the attachment of significance to this deviation should be viewed with caution.

Three degrees of freedom are associated with the specification of Earth's luminous flux in any kind of steady state: (I) the earth's phase angle from the direction of measurement; (2) the geographic longitude passing through that point on the Earth's surface where the direction of measurement passes through the zenith; and (3) the corresponding latitude. In the present study we may assume that the third degree of freedom has been held approximately constant. It is therefore of significance to ask what the locus of relative brightness extrema in the (Earth phase angle, sub-Mariner longitude) plane looks like. Table I has been examined for relative maxima and minima and the associated data have been plotted in Figure . There does not appear to be a very systematic effect associated with phase. The longitudes associated with the most pronounced maxima correspond to a point in the Atlantic Ocean above the eastern extreme of South America, and appear to correspond very nearly to a maximum in projected land-fraction. The most pronounced minima is in the center of the Pacific Ocean and represents approximately a minimum in projected land fraction.

In order to obtain an estimate of the relative reflecting power of landforms compared to ocean, the following simple analysis, which ignores any possible limb-darkening the Earth may exhibit, was undertaken. A globe of the Earth was photographed with the sub-camera point corresponding to the geographic coordinates of the sub-Mariner point for the means of each of the above two brightness extrema. The two photographs were planimetered over the entire projected disk. Thus the phase-angle was assumed zero. Table I was then examined for maximum-minimum pairs which satisfied the criteria that the maxima fall between 280° and 360° longitude and the minima fall between 170° and 240°. In addition the elements of a given pair were required to differ by less than 1.5 in phase angle. Twenty-two such pairs were found. The ratio of maximum to minimum

of this fact and the fact that no distinct trend is exhibited against phase angle in b.

Figure 7, the assumption of a full earth in planimetering the photographs seems justified.

The results of the planimetry yield that the projected land fraction for the Pacific point is 0.06l while that for the Atlantic sub-Mariner point is 0.355. The mean ratio of maximum to adjacent minimum brightness is $1.12 \pm .01$ S.D. Simple calculations then reveal that, for equal illuminated areas, average landform is $1.41 \pm .03$ S.D. times as bright as average ocean. Water has a higher albedo than land, but is not as good a back-scatterer. Inasmuch as the range in phase angle was not too far from zero, this might be part of the explanation for these relative reflectivities. However, one would thus expect to see the ratio of maximum to minimum brightness exhibit a phase dependence, contrary to the present observations. An alternative explanation would suggest that clouds, phenomena of high albedo, tend to concentrate over land relative to water.

Two exceptional cases are noted in Figure where maxima rather than minima occur over the Pacific. Meteorological reports from the vicinity at these times do not suggest exceptional cloudiness. Perhaps the ocean region from which a specular reflection of the sun is obtained was unusually becalmed. A more detailed mathematical analysis of these data entailing a running correlation with world-wide meteorological conditions and especially Tiros satellite photographs will be the subject of a later paper.

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REFERENCES

- Bergh, S. van den, The Color of the Moon, Astron. J., 67, 147-150, 1962.
- Danjon, A., Albedo, Color, and Polarization of the Earth, <u>The Earth as a Planet</u>, edited by G. P. Kuiper, University of Chicago Press, Chicago, 726–738, 1954.
- Johnson, H. L., The UBV Photometric System, Ann. d'Astrophys., 18, 292-296, 1955.
- Mathews, T. A., and A. R. Sandage, Optical Identification of 3C48, 3C196, and 3C286 with Stellar Objects, Astrophys. J., 138, 30–56, 1963.
- McLauchlin, J., Symposium on Spectroscopic Instrumentation in Extra-Terrestrial

 Environments, Jet Propulsion Laboratory, paper presented April 28, 1964.
- Michard, R., On the Energy Distribution in the Ultra-Violet Continuous Spectrum of the Sun, Bull. Astron. Inst. Netherlands, 11, 227-231, 1950.
- Minnaert, M., The Photosphere, <u>The Sun</u>, edited by G. P. Kuiper, University of Chicago Press, Chicago, 88–177, 1953.
- Radio Corporation of America, RCA Photosensitive Devices and Cathode Ray Tubes,
 Radio Corporation of America, Harrison, New Jersey, 1–31, 1958.
- Rougier, G., Photoelectric Color Index of the Moon, Ann. Obs. Strasbourg, 3, 257–260, 1937.
- Stebbins, J., and G. E. Kron, The Stellar Magnitude and Color Index of the Sun, Astrophys. J., 126, 266-28-, 1957.
- Walsh, J. W. T., Photometry, Constable and Company Ltd., London, Appendix IV, 1953.
- Wildey, R. L., and H. A. Pohn, Detailed Photoelectric Photometry of the Moon, in preparation, 1964.
- Wildey, R. L., and B. C. Murray, Ten Micron Stellar Photometry First Results and Future Prospects, Colloque International d'Astrophysique tenu a l'Universite de Liege,

 Tome IX, 460–468, 1964.

TABLE 1

7 .176	330	-1.9	36.7	1041	88.0	2.56
7 .500	2 13	-1 .9	36.3	1049	87 .7	2.43
7 .708	139	-1 .9	36.l	1055	87 .5	2.35
7.965	48	-1.9	35.7	1061	87 .3	2.48
8 .213	316	-1 .9	35.4	1069	87.0	2.66
8 .408	246	-1 .9	35.3	1073	86 7	2.52
8 . 4 79	220	-1.9	35 .1	1075	86.3 86.0	2.41
8 .602	175	-1 .9	35.0	1078	86.3	2.31
8 .876	76	-1 .9	34.6	1086		2.46
9.194	321	-1 .9	34.3	1095	8 6 .7	2.63
9.362	260	-1 .9	34.l	1099	87.0	
9.455	227	-1.9 -1.9	34.0	1101	87 .3	2.53 2.44
9.559	189	-l .9	33 .9	1104	87 .7	2.34
9.686	144	-l .9	33 .7	1108	88 .0	2.49
9.885	71	-1.9	33.3	1115	87.9	2.64
10.100	354	-1 .9	33.l	1119	87 .8	2.77
10.166	330	-1 .9	33.0	1121	87 .7	2.96
10.220	310	-1.9	32.9	1122	87 .6	2.79
10 .304	280	-1 .9	32.8	1124	87 .4	2.68
10 .362	260	-1 .9	32.7	1126	87 .3	2.55
10.538	196	-1.9	32.6	1133	87.2	2.46
10.788	106	-1 .9	32.3	1138	87 .0	2.60
10.949	47	-1 .9	32.1	1142	87 .0	2.74
11 .174	326	-l .9	31 .8	1150	87 .0	2.90
11 .279	288	-l .9	31 .7	1152	87 .0	2.79
11.293	284	-1.9	31 .7	1153	87.0	2.68
II .592	175	-1 .9	31 .3	1162	87 .0	2.59
II .866	75	-1 .9	31 .0	1170	87 .0	2.62
11.907	61	-1.9	30.9	1171	87.0	2.88
11.988	32	-1 .9	30.8	1173	87 .0	2.77
12.126	342	-1 .9	30.7	1177	87 .0	2.91
12.708	132	-1.9	30.0	1195	87 .0 87 .0	2.74
13 .130	339	-1 .9	29 . 4	1205	86 7	2.88
13 .358	257	-1 .9	29 . l	1213		2.81
l3 .971	35	-1 .9	28 .3	1231	87 .5	2.90
l4 .407	238	-1 .9	27 .9	1243	86 .3	2.79
14.465	217	-1.9	27.8	1246	85.8	2.69
14 .538	190	-l .9	27 .7	1248	85 .5	2.79
14 .956	39	-l .95	27 .1	1261	85 .4	2.92
15 .347	259	-1.96	26.7	1272	85.3	2.90
l5 .443	224	-1 .96	26.6	1274	85.2	2.79
l5 .54l	187	-1 .97	26. 4	1279	85.0	2.91
15 .871	68	-1 .98	26.0	1288	85 .0	3.12
	298	-1 .99	25.6	1300	85 .0	3.03
16 .230 16 .431	226	-1 .99	25.4	1305	85 .0	2.92
16 .604	164	-2.00	25 . 2	1311	85.0	3 .07
16 .678	137	-2.00	25 . l	1314	85.0	3 .25
l6 .7 6 7	105	-2.00	25.0	1316	85.0	3.08
l6 .824	85	-2.00	24.9	1318	85 0	2.94
l6 .973	30	-2.0l	24.8	1323	85.0	3.13
17 .137	331	-2.02	24.6	1328	85.0	3.15
17 .184	314	-2 .03	24.5	1330	85.0	3 .32
17 .211	304	-2 .04	24.4	1331	85.0	3 .16
17 .783	99	-2.05	23 .8	1348	85.0	3 .10
18 .356	251	-2.05	23 .1	1368	85.0	3 .07
U	ZJI	2.00	20.1	1300	05.0	J .07

18.428	225	-2.06	23.0	1371	85.0	2.93
18 .505 18 .881	196 61	-2.07 -2.08	22.9 22.5	1373 1386	85 .0 85 .0	3 .09 3 .30
19.263	282	-2.11	22.0	1398	85.0	3.17
19 .395	220	-2.11	21.8	1402	85.0	3.16
20.132	329	-2.11	21.0	1427	84.9	3.17
20 .444 20 .673	216 132	-2.l -2.2	20.7 20.4	1437 1446	84.7 84.4	3.20 3.39
21 .090	343	-2.2	20.4	1461	84.0	3.39
21 .458	210	-2.2	19.6	1473	84.0	3.26
21 .833	74	-2.2	19.1	1487	84.0	3.41
26 .301	26l	- 2.50	15.5	1616	83.0	3.52
26 .493 26 .697	192 118	-2 .52 -2 .53	15.4 15.2	1625 1632	82.5 82.0	3 .30 3 .34
26 .966	20	-2 .56	15.0	1636	82.0	3.34
27 .068	284	-2.59	14.8	1654	82.0	3.67
27 .399	224	-2.62	14.7	1660	82.0	3.44
27 .823 28 .073	71 341	-2.66	14.3	1677	82.0 82.0	3.38
28 .073 28 .198	295	-2 .67 -2 .68	14.1 14.0	1686 1693	82.0	3 .55 3 .82
28.241	279	-2.68	13.9	1698	82.0	3.60
28 .467	199	−2 .71	13.8	1704	82.0	3.51
28.698	116	-2.72	13.6	1714	82.0	3 .41
28 .948 29 .097	24	-2.75 -2.79	13 .4 13 .3	1724 1734	82.0 82.0	3.60
29.333	302 245	-2 .78 -2 .79	13.3	1734	82.0	3.73 3.64
29.435	209	-2.80	13.1	1744	82.0	3.53
29.687	117	-2 .83	12.9	1756	82.0	3.43
29.908	22	-2.85	12.7	1767	82.0	3.62
30 .132 30 .305	317 253	-2 .80 -2 .89	12.6 12.5	1775 1783	82.0 82.0	3.80
30.694	113	-2.87 -2.94	12.4	1800	82.0	3 .68 3 .6l
31 .111	321	- 2.97	12.2	1818	82.0	3 .83
31 .257	269	-2 .98	12.1	1826	82.2	3 .70
31 .406	216	-2.99	12.1	1830	82.5	3 .63
31 .492 31 .543	186 171	-3 .01 -3 .03	12.0 11.9	1835 1838	82 <i>.</i> 8 83.0	3 .51 3 .64
31 .630	136	-3 .05	11.9	1881	82.8	3.75
31 .811	71	-3 .07	11.8	1850	82.5	3.72
32.032	350	-3.09	11.8	1860	82.3	3 .83
32.201 32.281	288	-3 .23 -3 .24	.4 .4	1915 1918	82.0 82.7	3.85
32.785	259 77	-3 .24 -3 .32	11.4	1942	83.3	3.63 3.49
33 .295	252	-3.38	11.4	1966	84.0	3.41
33 .471	189	-3 .4l	11.3	1977	84.0	3.25
33 .641	127	-3 .5	11.3	1986	84.0	3.49
33 .883 34 .214	39 280	-3 .5 -3 .5	II .3 II .3	1997 2014	84.l 84.2	3 .35 3 .49
34.333	238	-3.5 -3.5	11.3	2020	84.3	3.47
34 .431	203	-3.5	11.3	2024	84.4	3.17
34 .525	167	- 3.5	11.3	2029	84.5	3 .37
34 .636	127	-3.5	11.3	2034	84.6	3.49
34 .753 34 .818	86 64	-3.5 -3.5	11 .4 11 .4	2040 2044	84.6 84 <i>.</i> 7	3.42 3.24
34 .894	34	-3.5 -3.5	11.4	2048	84.7	3.45
35 .031	345	-3.5	11.4	2055	84.8	3.57

35 .169	296	-3 .6	ll .4	2063		3.50
35 .239	270	-3 .7	ll .4	2066		3.32
35 .314	243	-3 .7	II .4	2070	85.0	3.23
35 .434	201	-3 .8	II .4	2076	85.2	3.05
35 .533	164	-3.8	II .5	2082	85.4	3.28
35 .639	125	-3.8	II .5	2089	85.5	3.39
35 .730	92 62	-3 .8 -3 .8	II .5 II .6	2093 2097	85.6 85.6	3.32
35 8l5 35 .909	28	-3 .8	11.6	2100	85.7	3.14 3.36
36 .028	345	-3 .8	II .6	2107	85 .7	3.44
36 .192	285	-3 .8	II .7	2117	85 .8	3.52
36 .265	259	-3 .8	11 .7	2120	85.9	3.4l
36 .385	21 6	-3 .9	11 .7	2126	86.0	3.25
36 .739	88	-3 .9	11.9	2140	86.0	3.29
37 .042	339	-3 .9	12.0	2160	86.5	3.58
38 .107 38 .410	314 206	-4.0 -4.1	12.5	2220 2230	87 .0 88 .0	3.34
38 .773	76	-4.2	12.8	2250	89.0	3.30 3.15
38 .850	45	-4.2	12.9	2260	89.5	3 .31
39 .044	336	-4.2	13.0	2270	90.0	3 .33
39 .244	263	-4 .2	13 .1	2280	90.0	3.39
39 .398	244	-4 .2	13 .2	2290	90.5	3.32
39 .352	225	-4.2	13 .2	2290	91 .0	3.22
39 .447	191	-4.2	13 .3	2300	91 .5	3.14
39 .556	149	-4 .3	13.4	2300	92.0	3.25
39 .627	125	-4 .3	13.4	2310	92.2	
39 .684	105	-4.3	13.4	2310	92.4	3.32
39 .750	81	-4.4	13.4	2320	92.6	
39.848	45	-4.4	13.5	2320	92.8	3 .24 3 .35
39 . 979	358	-4 .4	13.6	2330	93.0	3 .53
41 .153	294	-4 .5	14.4	2400	93.0	3 .17
41 .185	281	-4.6	14.4	2400	93 .5	3.07
41 .441	1 <i>9</i> 0	- 4.6	14.5	2410	94 .0	2.95
4l .760	74	-4 .7	14.8	2440	95 .0	2.88
4l .845	43	-4 .7	14.9	2450	95 .3	2.94
41 .935	ii	-4.7	14.9	2450	95.7	3.II
42.031	342	-4.8	15.0	2460	96.0	3.05
42.310	236	-4.8	15.0 15.2 15.4	2470 2490	96.5 97.0	3.02
42.579 42.644	138 115	-4.8 -4.8	15.5	2500	97 .3	2.83
42.715	90	-4.9	15.5	2500	97 .5	2.78
42.752	76	-4.9	15.6	2510	97 .8	2.70
42.799	60	-4.9	15 .6	2510	98.0	2.79
42.989	350	-4.9	15 .8	2520	98.2	2.89
43 .192	276	-5 .0	15.9	2530	98.4	2.87
43 .382	208	-5 .0	16.0	2540	98.6	2.78
43 .553	146	-5 .0	16.2	2550	98 .8	2.70
43 .763	70	-5 .1	16.3	2570	99 .0	2.69
44 .007	345	-5 .l	16.5	2590	99.5	2.83
44 .075	3I8	-5 .l	16.5	2590	100.0	
44 .229 44 .371	262	-5.2 -5.2	16.6 16.8	2600 2610	100.5	2.80
44 .506	211 162	- 5.2	16.9	2620 2630	101.0 101.5 102.0	2.69 2.63
44 .647	112	-5 .2	17.0	2000	102.0	2.58

44.726	83	-5.3	17.1	2630	102.1	2.37
44.807	54	-5.3	17.2	2640	102.2	2.60
45.011	341	- 5 .3	17.3	2650	102.4	2.76
45 .218	264	-5.4	17.4	2670	102.5	2.69
45.328	225	-5.4	17.6	2680	102.7	2.51
45.845	39	- 5 .5	0.81	2710	102.8	2.37
46.277	238	- 5.6	18.3	2750	0. 201	2.34
46.496	160	- 5. 6	18.5	2760	104.5	2.36
4 6 .730	80	- 5 .7	18.6	2770	0. 601	2.34
47.032	330	- 5 .7	19.0	2800	106.3	2.51
47.326	225	- 5 .7	19.1	2820	106.7	2.44
47.399	198	- 5.8	19.2	2820	107.0	2.36
47.465	174	- 5 .8	19.3	2830	107.3	2.31
47.569	136	- 5 .9	19.3	2830	107.7	2.27
47.715	84	- 5 .9	19.5	2840	0.801	2.23
47.780	6l	-6.0	19.5	2850	108.3	2.22
48.081	312	-6.0	19.8	2870	6.80	2.41
48.459	174	-6 .0	20.1	2910	8.801	2.27
48.518	153	-6 .[20.1	2910	109.0	2.18
48.666	100	-6.2	20.3	2920	109.3	2.30
48.986	345	-6.2	20.6	2940	109.7	2.38
49.197	268	- 6.2	20.7	2960	110.0	2.61
49.232	256	-6 .2	20.8	2960	110.5	2.59
49.309	228	-6.2	20.8	2970	111.0 111.5	2.56
49.458	174	-6.3	20.9	2980	III .5 III .5	2.31
49.527	149	-6.3	21.0	2980 2980	111.7	2.32
49.572	133	-6.3	21 .0 21 .1	2990 2990	111.7	2.26
49 .638 49 .708	109 94	-6.3 -6.3	21 .1 21 .1	3000	111.8	2.23
49.878	23	-6.4	21.1	3010	112.0	2.32
50.086	309	-6.4 -6.4	21 .3	3020	112.2	2.42
50.159	282	-6.4 -6.4	21 .4	3030	112.3	2.54
50 .187	282 271	-6.5	21.5	3030	112.4	2.57
50 .275	239	-6 .5	21.6	3 0 30	112.6	2.69 2.49
50 .397	196	-6.5	21 .0 21 .7	3040	112.7	2.49
50 .461	172	-6.5	21.8	3060	112.8	2.32
50 .534	146	-6.5	21.8	3060	112.9	2.32
50 .635	iii	-6.6	21.9	3070	113.0	2.36
50 .739	72	-6.6	21.9	3080	113 .5	2.49
51 .032	328	-6 7	22.1	3100	114.0	2.71
51 .288	235	-6.7	22.4	3120	114.5	2.64
51 .318	224	-6.7	22.4	3130	115.0	2.56
51 .600	121	-6.7	22.7	3140	115.2	2.54
52.052	319	-6.8	22.9	3180	115 .5	2.80
52.281	239	-6.9	23.0	3210	115.7	2.63
52.340	213	-6.9	23 .1	3210	116.0	2.53
52.399	192	-7 .0	23.2	3210	116.3	2.51
52 . 458	171	-7 .0	23.4	3220	116.5	2.53
52.618	113	- 7 .0	23 .5	3230	116.7	2.60
53.000	336	-7 .1	23.7	3270	117.0	2.83
53.254	245	-7.2	23.9	3290	117.5	2.73
53 .288	233	- 7.2	24.0	3290	118.0	2.71
53 .538	142	- 7.2	24.1	3310	118.5	2.74
53 .817	26	-7 .3	24.4	3340	119.0	2.93